



UGC 5340



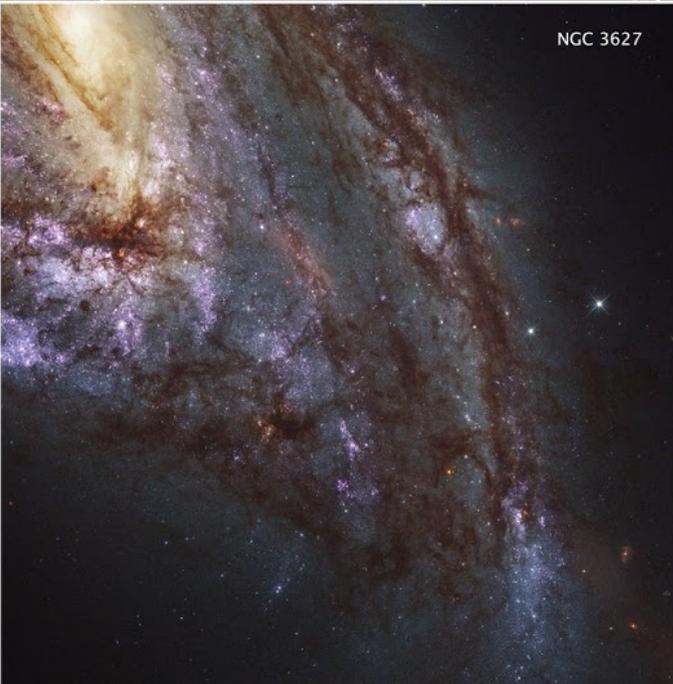
NGC 4258



UGCA 281



NGC 3368



NGC 3627



NGC 6744

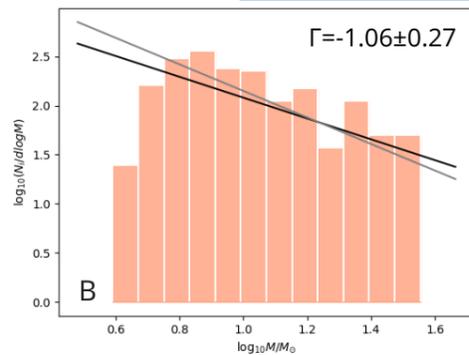
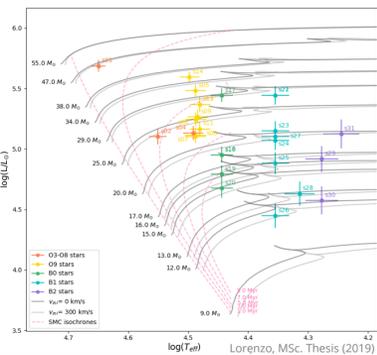


M. Lorenzo: OSIRIS@GTC

metal poor galaxies, large sample OB-stars → regions of intense SF show flatter IMF

(Lorenzo, MSc Thesis 2019)

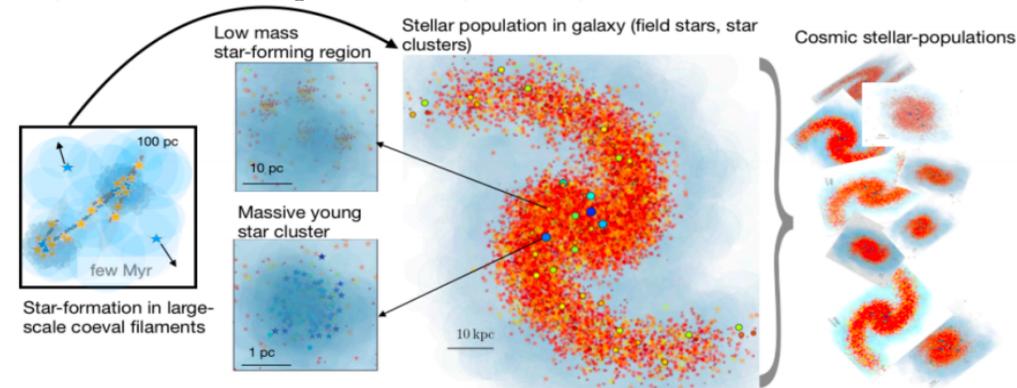
Our team is working to make Sextans A the new low-Z standard



T. Jerabkova: GAIMF →

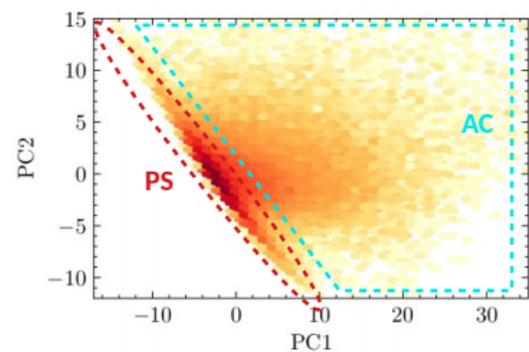
IGIMF connects small-scale SF to galaxy-wide properties.

The IGIMF theory uses empirically driven constraints to compute stellar populations in star clusters, galaxies and on cosmic scales.

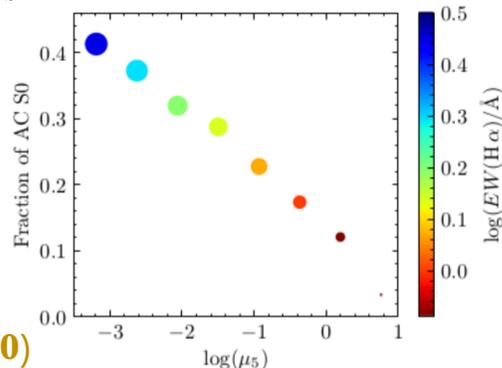


J.L. Tous: S0 subpopulations → PS (3/4) vs AC(1/4)

AC less massive, more luminous, with less concentrated light profiles, have a younger, bluer, and metal-poorer stellar component, and avoid high-galaxy-density regions



(Tous+20)

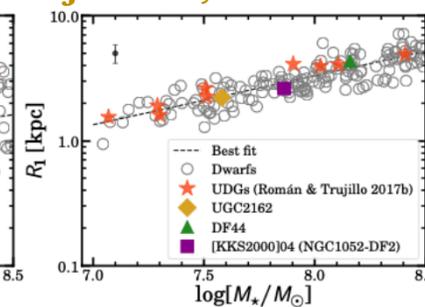
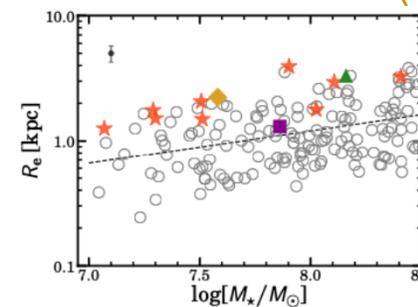
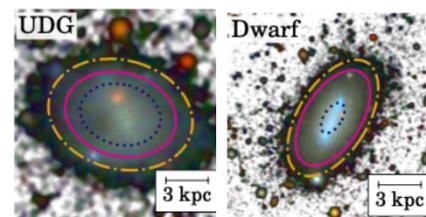


N. Chamba: SDSS Stripe82 survey

new physically motivated size R1 → expected location of the gas density threshold for star formation
UDGs are as small as dwarfs, not like the MW



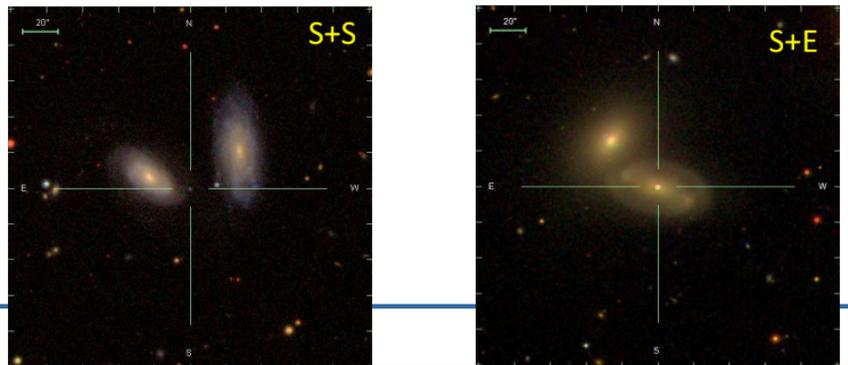
(Trujillo+20, Chamba+20)



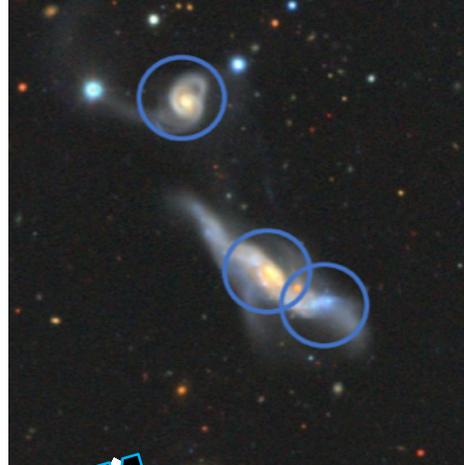
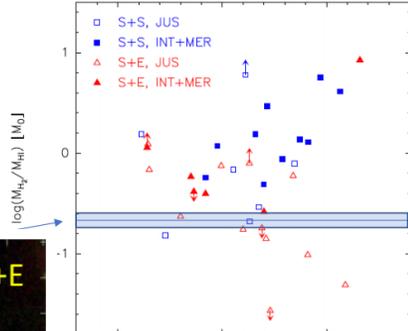
The nearby Universe: IMF, UDGs and Lenticulars

U. Lisenfeld: KPAIR

Pairs at different interaction phase
SF enhancement mostly for S+S pairs in first encounter, due to formation of M_{H2} from M_{HI}

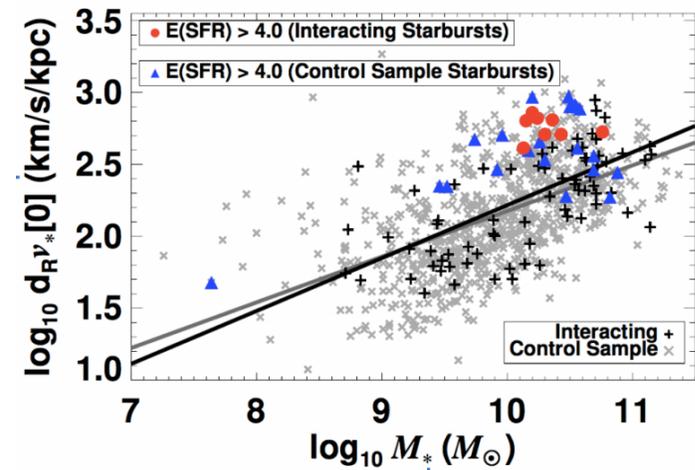
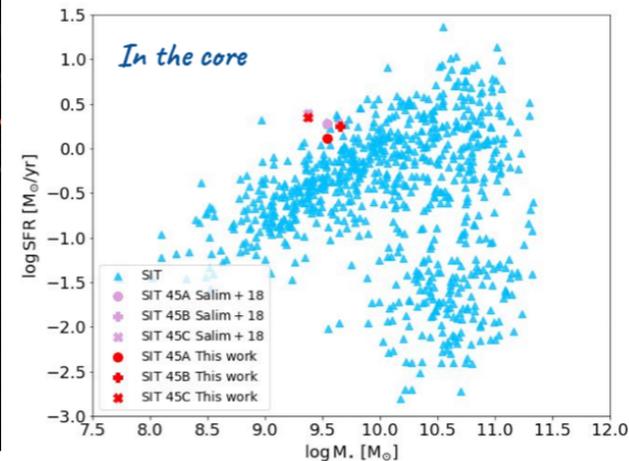


(Lisenfeld+19)



M. Argudo-Fernandez: SIT

SIT45 compact system (starburst galaxies triplet) → SFH shows it is an ongoing merger
(Argudo-Fernandez+15)

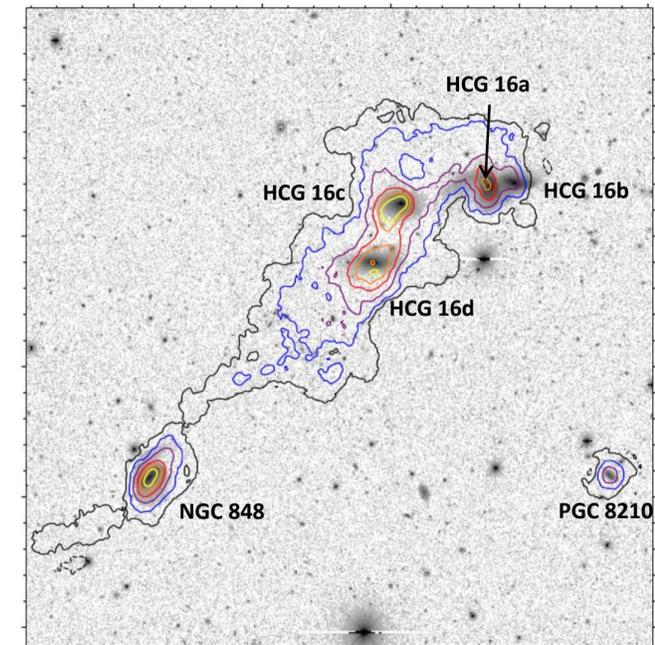


S. Diaz-Garcia:

mergers produce x2 enhancement in SF efficiency → Starbursts have higher central stellar concentrations, low gas depletion timescales, yet similar gas fractions as normal main-sequence galaxies
(Diaz-Garcia&Knapen20)

M. Jones: FAIR

Neutral gas study of HCG16 compact group but using FAIR (Findable, Accessible, Interoperable and Reusable) workflow for reproducibility
(Jones+19)

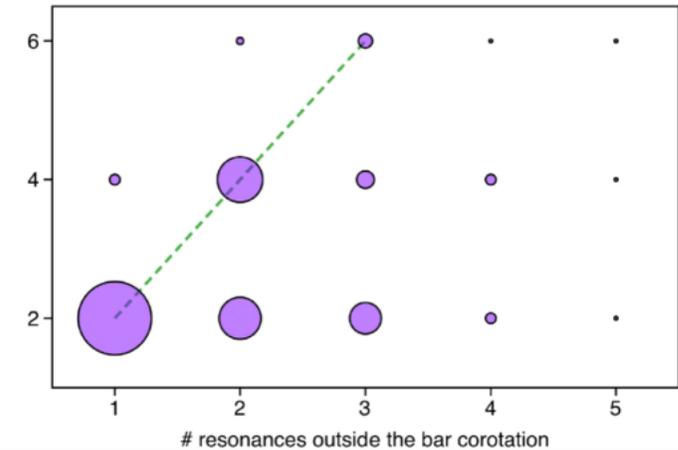
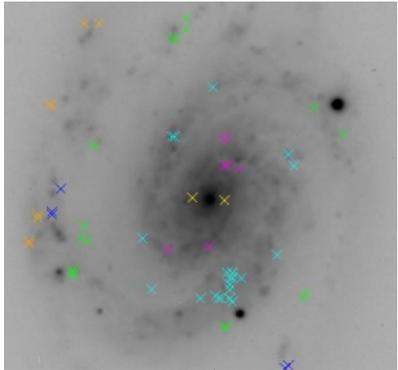


The nearby Universe: Mergers & interactions

J. Beckman: resonance patterns in discs of spiral galaxies is a two-dimensional pattern

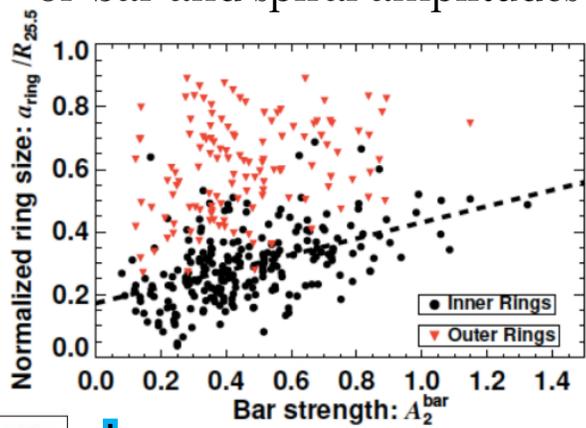
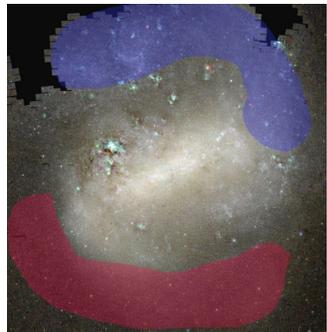
the number of the corotation detections increases as the number of arms grow

(Font+19)



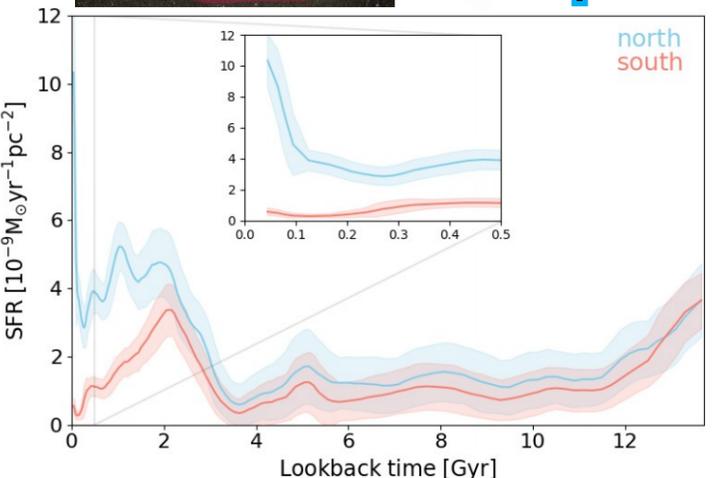
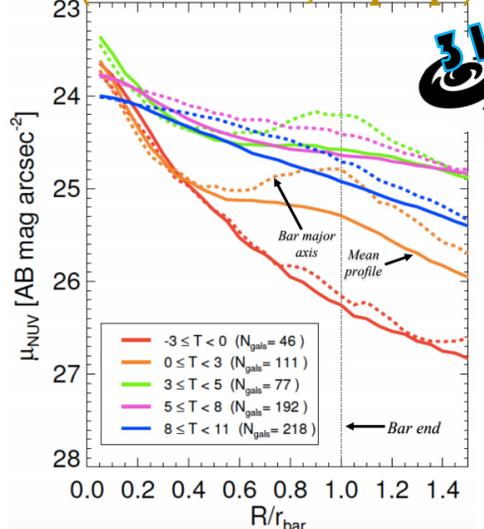
S. Diaz-Garcia: S4G
Bar strength correlates inner ring formation, coupling of bar and spiral amplitudes

(Ruiz-Lara+20)



(Diaz-Garcia+19a,b)

3 modes of distributed ionized gas within bars
(Diaz-Garcia, in prep.)



T. Ruiz-Lara:

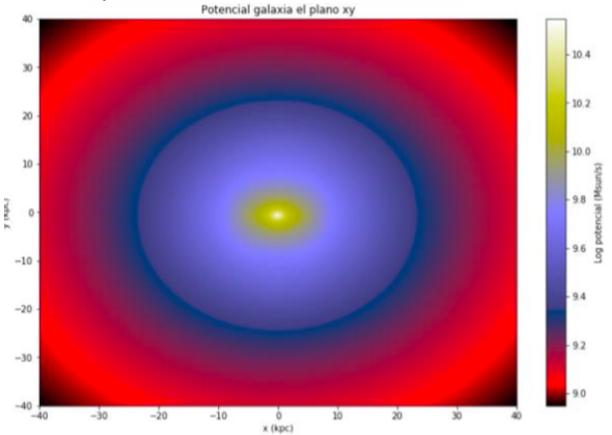
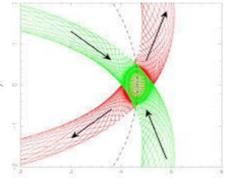
SMASH 2D photometric catalogues

formation telltale of the LMC spiral arm
clear dichotomy in the young stellar content of the two regions



V. Padura: AGAMA

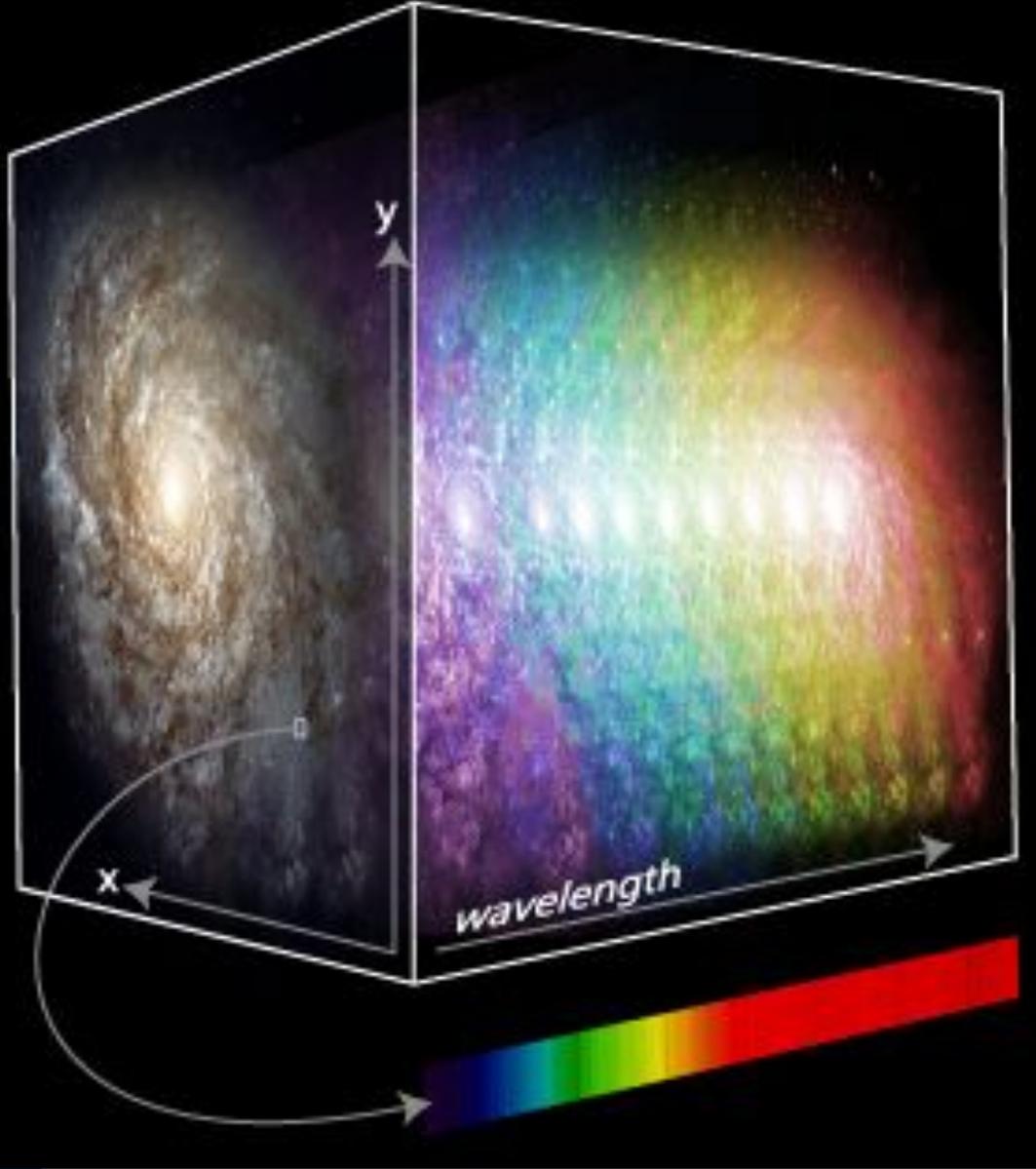
Testing N-body simulations to see if spiral arms can be triggered by unvariant manifolds



The nearby Universe: Arms & Bars

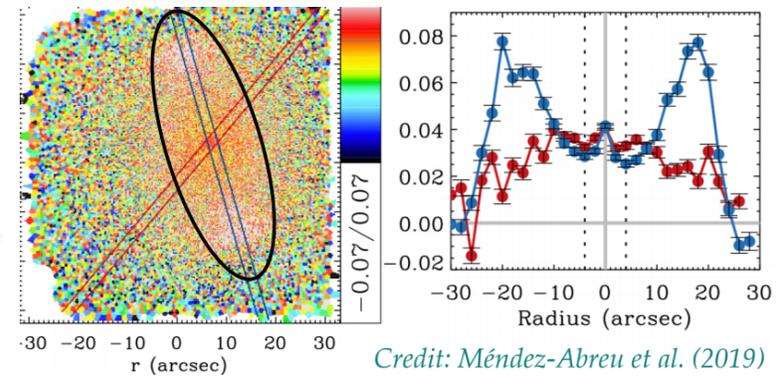
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A. De Lorenzo-Caceres:

first detection of a peanut structure in the inner bar of a double-barred system → long-lived structures, analogous to regular bars but with smaller size



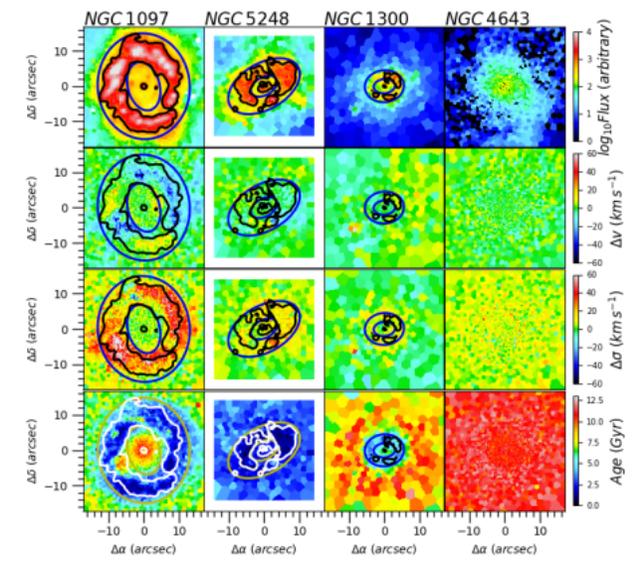
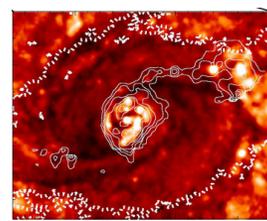
Credit: Méndez-Abreu et al. (2019)

(de Lorenzo-Cáceres+19)



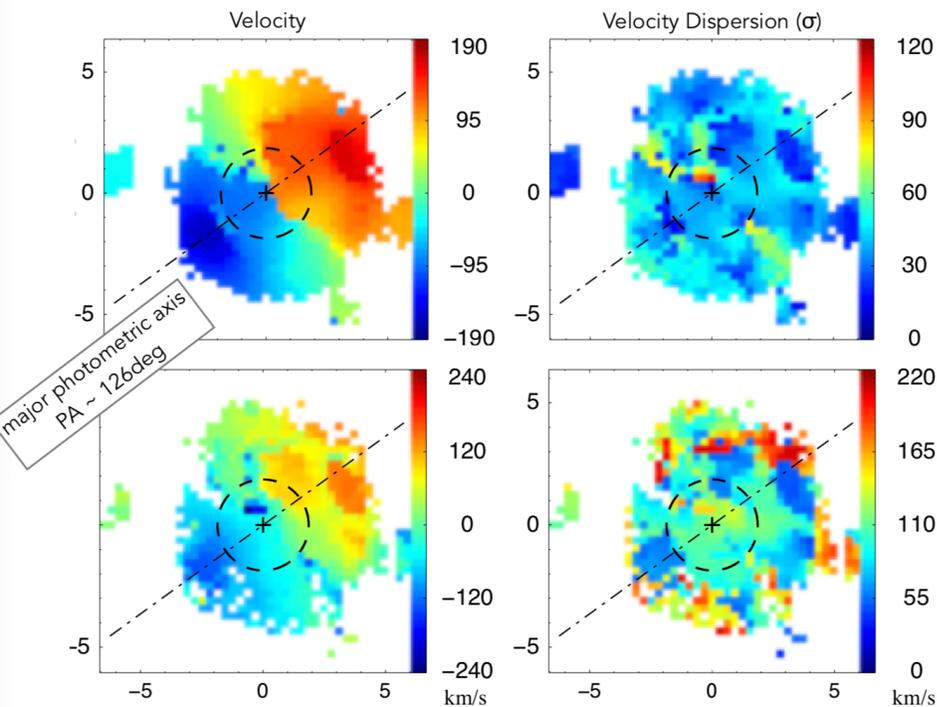
D. Rosado-Belza:

Stellar populations of galaxies with nuclear rings: Young stellar populations (mostly in nuclear rings) have higher velocities and velocity dispersion than the old stellar population.

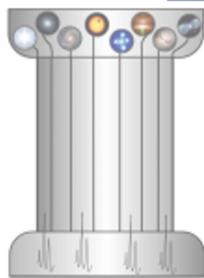


H α -[NII] emission in NGC7469 \rightarrow

presence of a very thin (20 pc) ionised gas disc supported by rotation, embedded in a thicker (222 pc), dynamically hotter one in NGC7469.



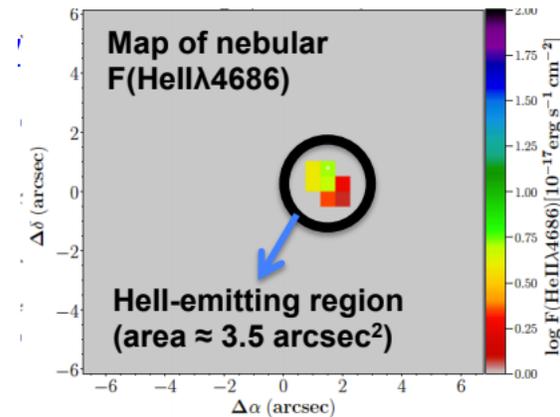
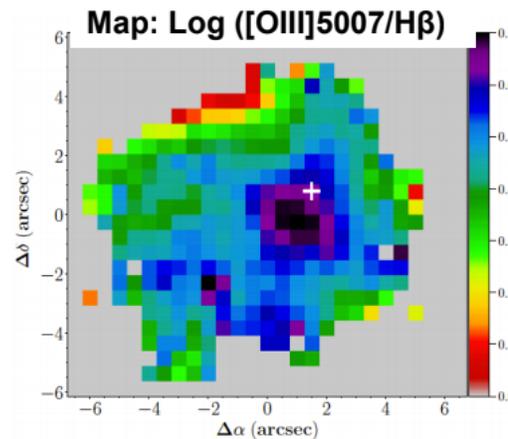
(Cazzoli+20)



Multi
Espectrógrafo en
GTTC de
Alta
Resolución para
Astronomía

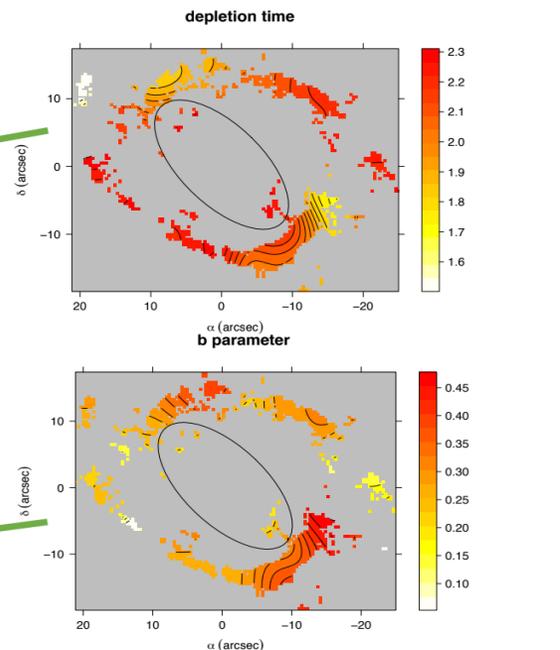
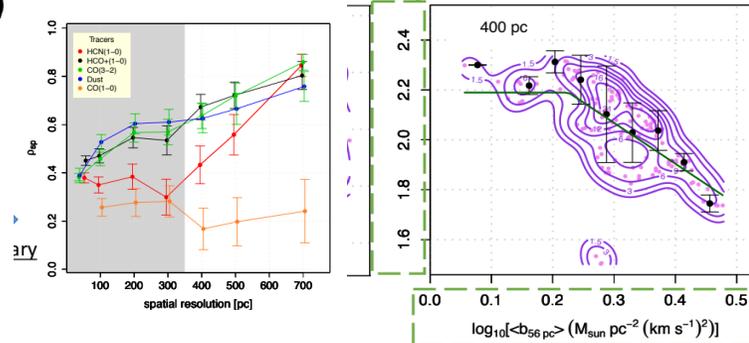
metal poor, high ionized HII PHL293B \rightarrow
analog of the early Universe

(Kehrig+19)



SFR in NGC1068 \rightarrow critical scale (300-400pc)

Two dynamical environments \rightarrow galactic dynamics key in deriving SF efficiency



The nearby Universe: IFS studies

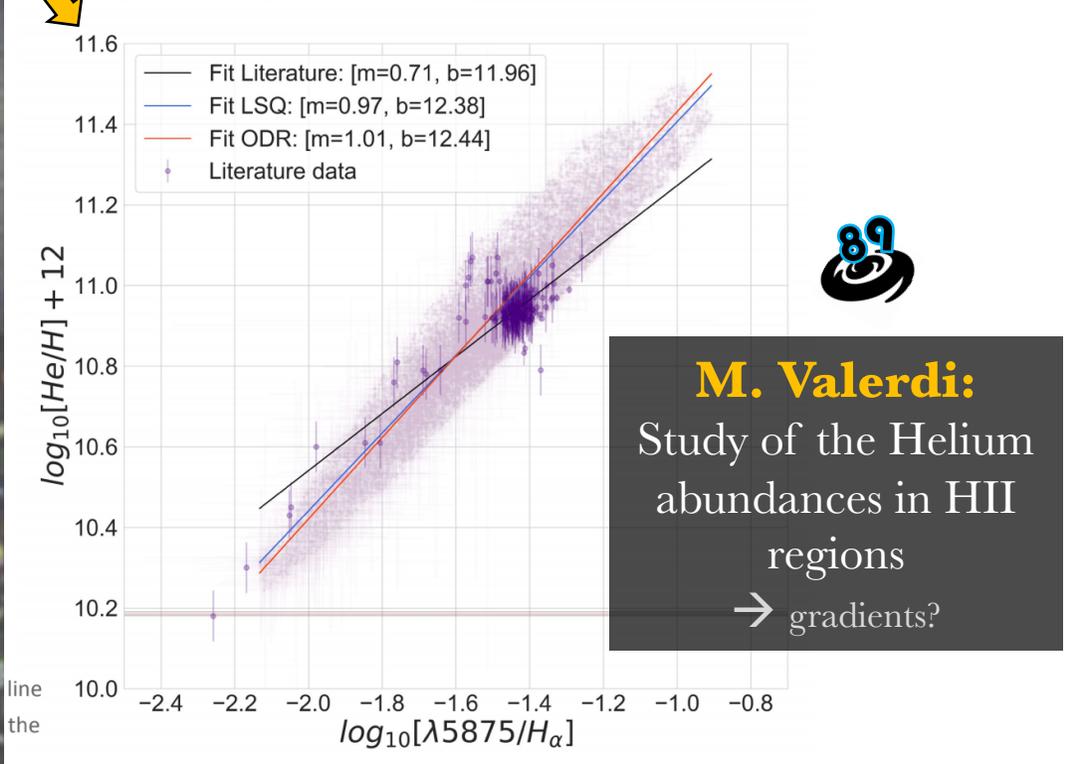
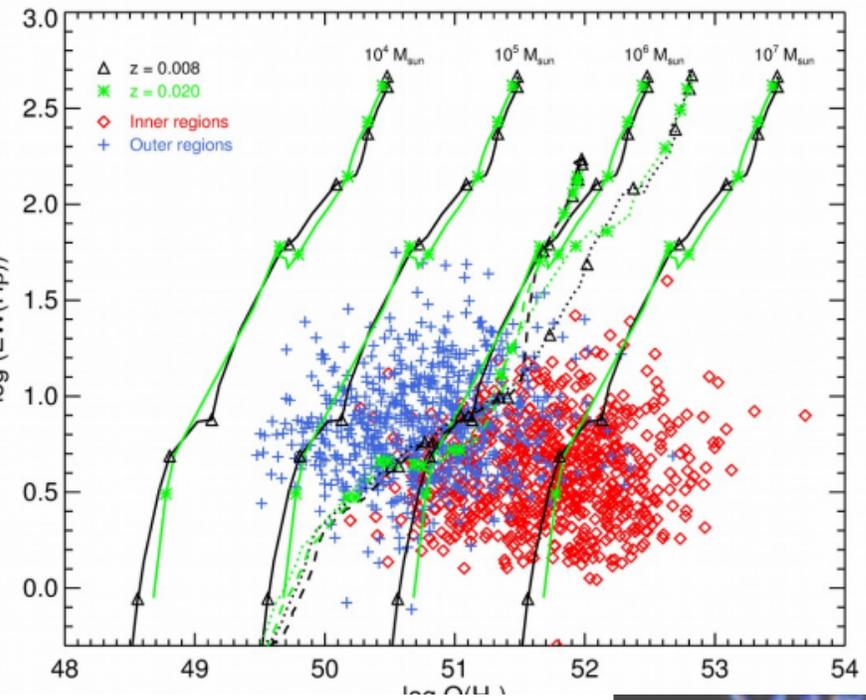
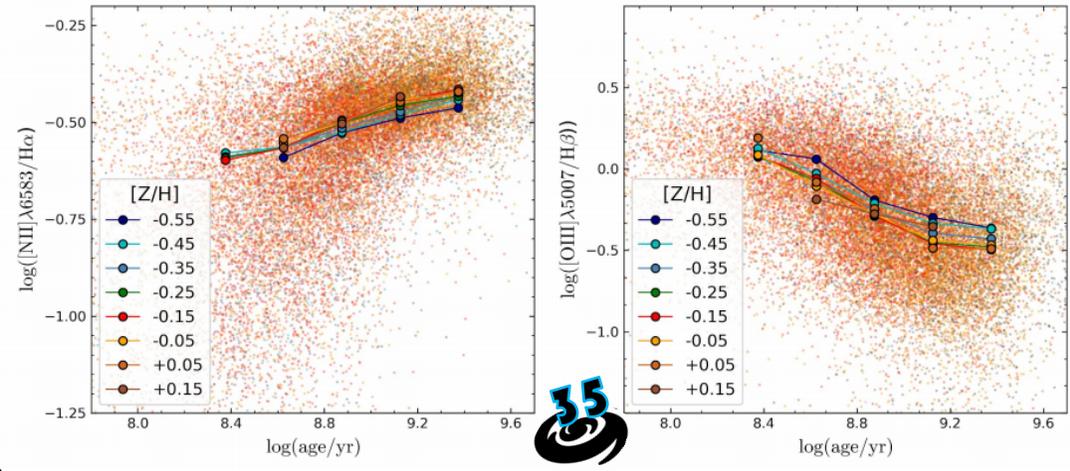
M. Rodriguez-Baras:

inner and outer HII regions show clear differences in metallicity, mass and age: Inner have higher metallicities, are older higher masses and larger sizes (Rodriguez-Baras+18,+19)



C. Espinosa-Ponce:

largest catalog of HII regions → new diagnostic diagrams → clear correlations of line ratios with both the ages and metallicities (Espinosa-Ponce+20)



M. Valerdi:
Study of the Helium abundances in HII regions → gradients?



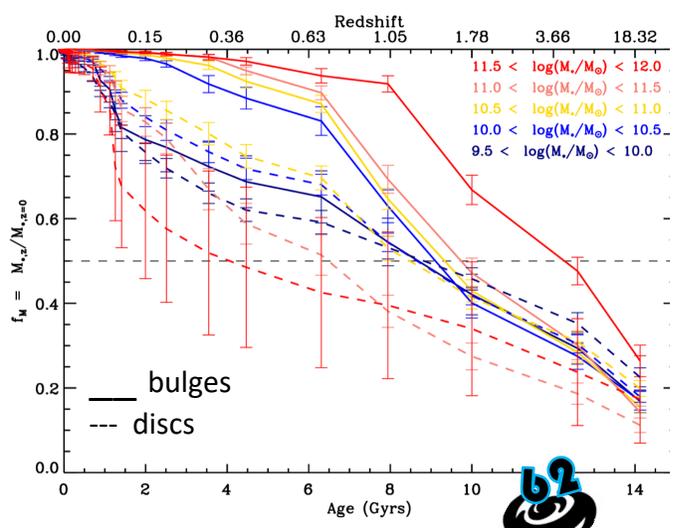
The nearby Universe: IFS studies

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Scaling relations

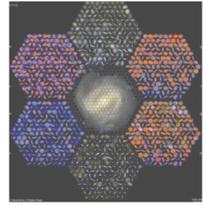


J. Mendez-Abreu:

Mass growth
inside-out scenario for non-barred disc galaxies

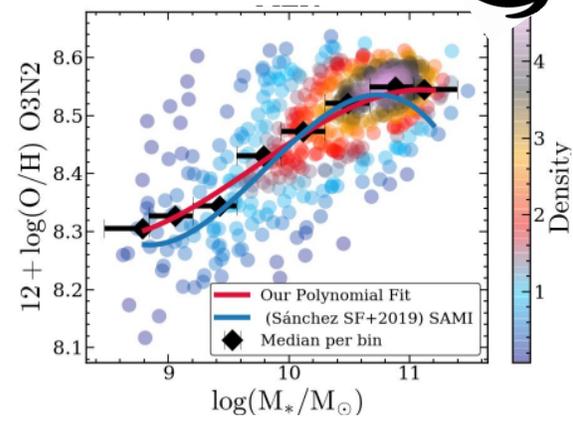
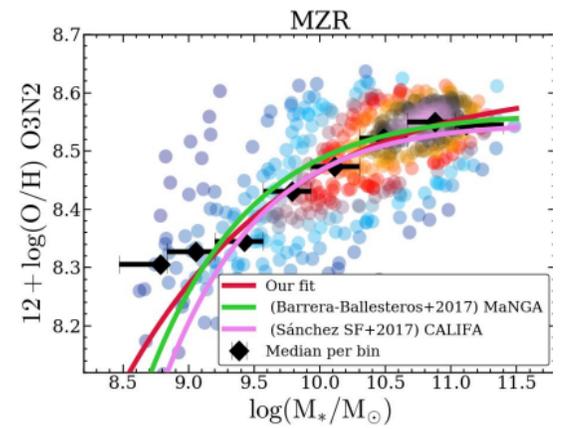
Bulges are older and more metal-rich and form earlier than discs

(Mendez-Abreu+19)



P. Alvarez-Hurtado: MZR

its functional form depends on statistical criteria → Generate non-existing secondary realtions! Careful!



The nearby Universe: IFS studies

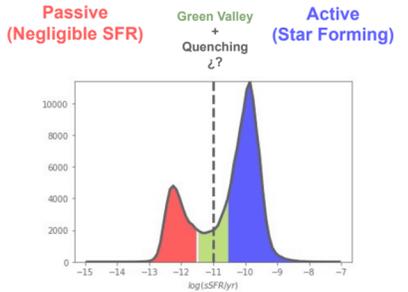
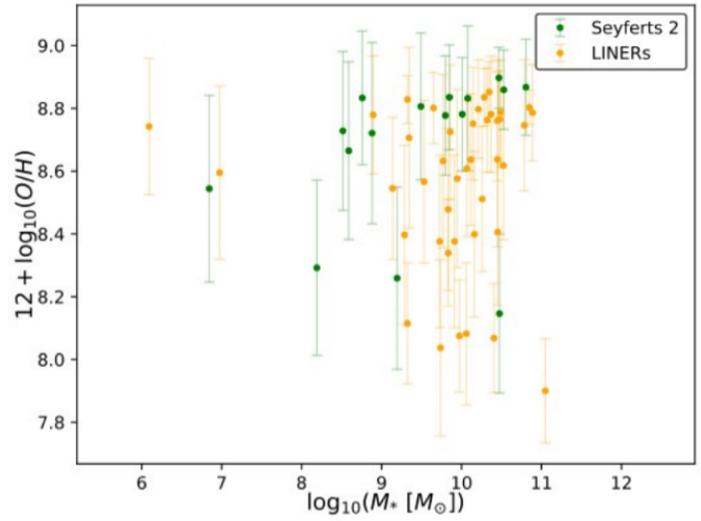
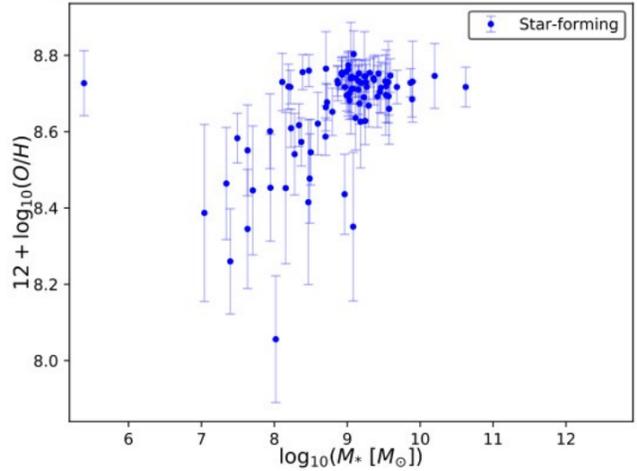
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B. Perez: Palomar Spectroscopic Survey

HII-CHI_MISTRY code → Metallicity relation

11 for the nuclear regions of galaxies

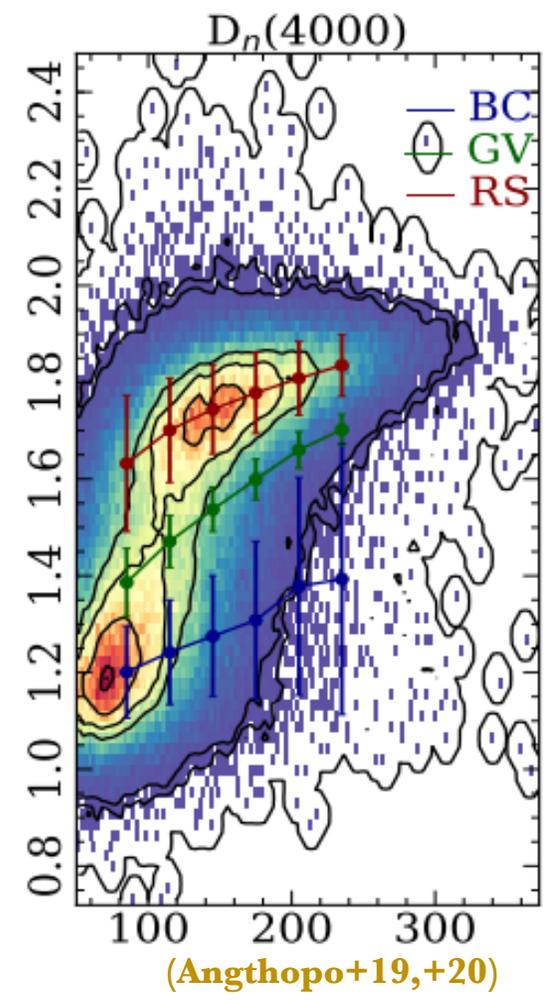


J. Angthopo: SDSS

Green Valley galaxies using

03 D4000 break → dust resilient

Quicker transition from BC to GV and slower from GV to RS

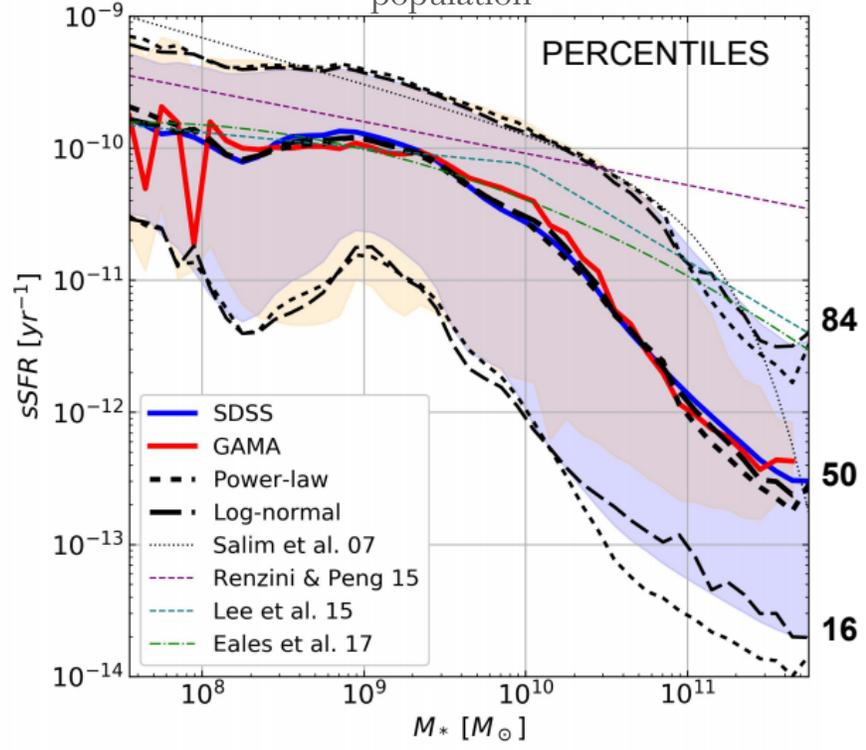


P. Corcho-Caballero: SDSS+GAMA



No bimodality →

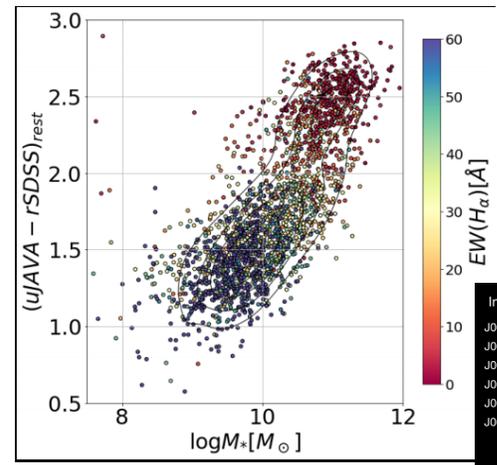
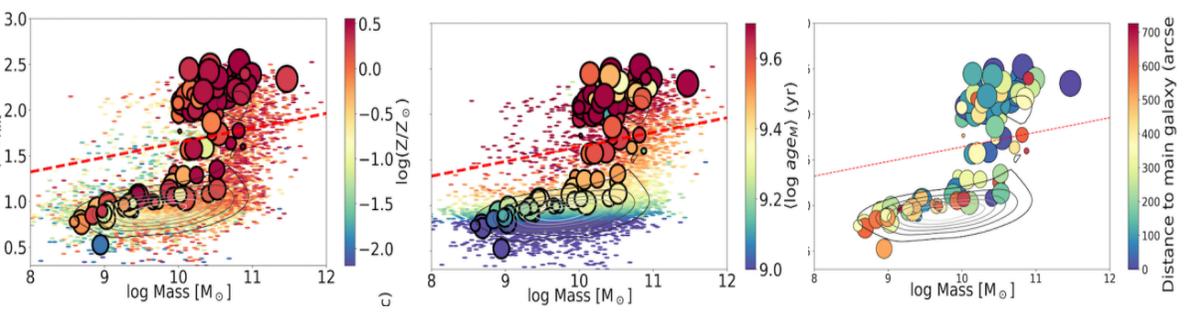
local Universe population has a single probability distribution + shallow power-law tails for the passive population



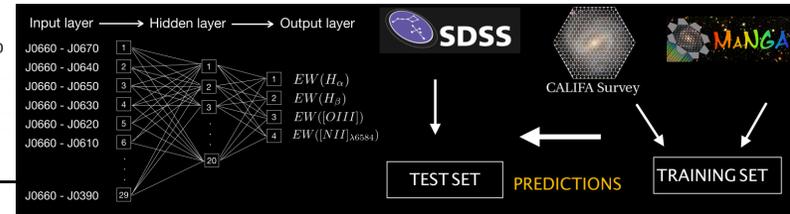
The nearby Universe: Global Scaling Relations



Cluster 1001 environment study of stellar populations → red galaxies are the closest to the BCG, most of the galaxies high t_{form} but low tau for RS

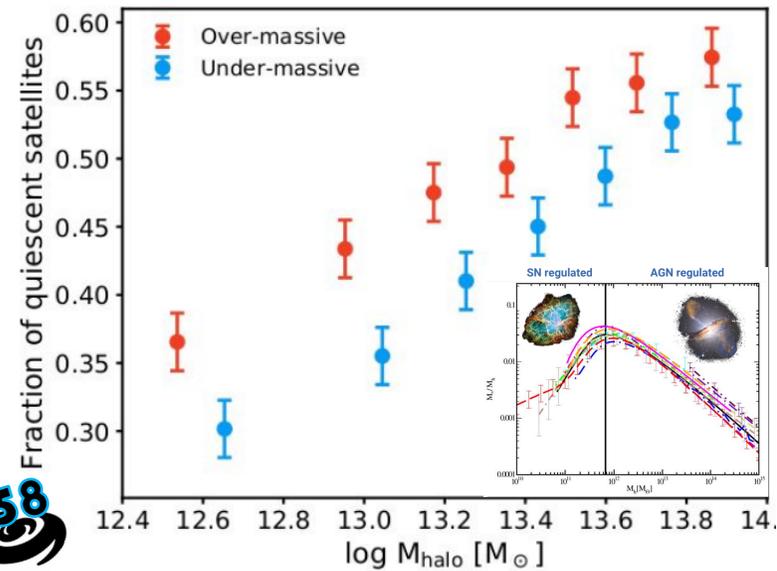


New method ANN (Artificial Neural Network) → Classify galaxies, measure emission lines, BPT diagrams, and study the luminosity function



I. Martin-Navarro:

BH feedback rules → halos hosting more massive black holes have an enhanced population of quiescent satellite galaxies and a hotter intracluster medium

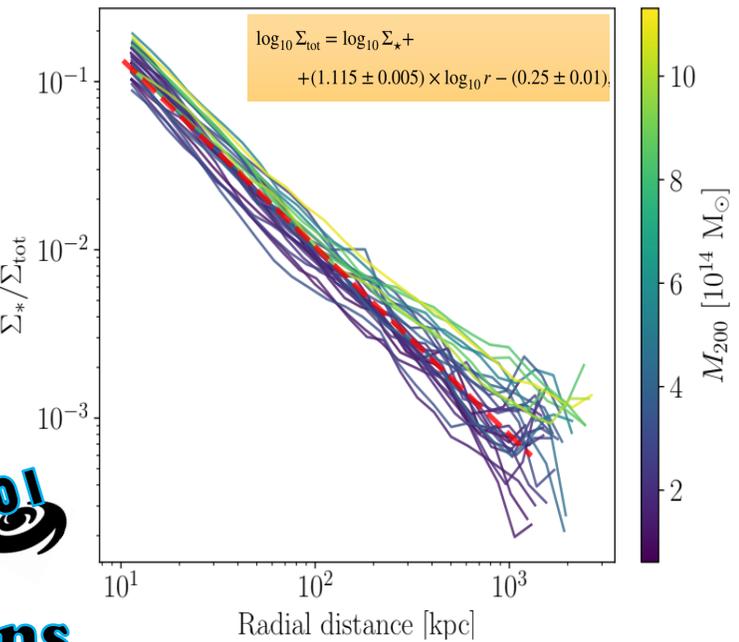


I. Alonso Asensio:

C-EAGLE Simulation

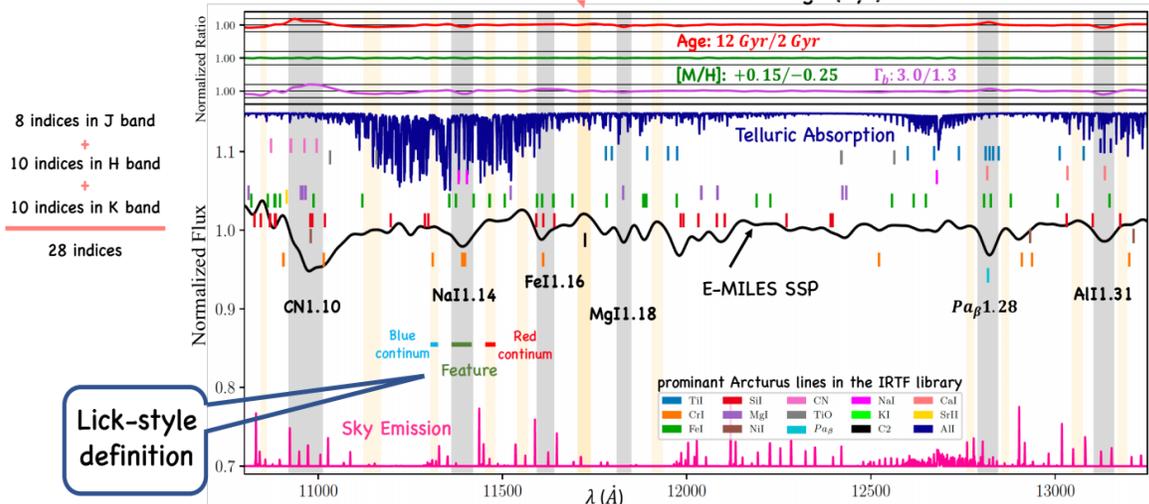
method to calculate halo mass and mass density from the ICL stellar mass density →

shape of the stellar mass distribution follows that of the total matter even more closely than observed, although their radial profiles differ



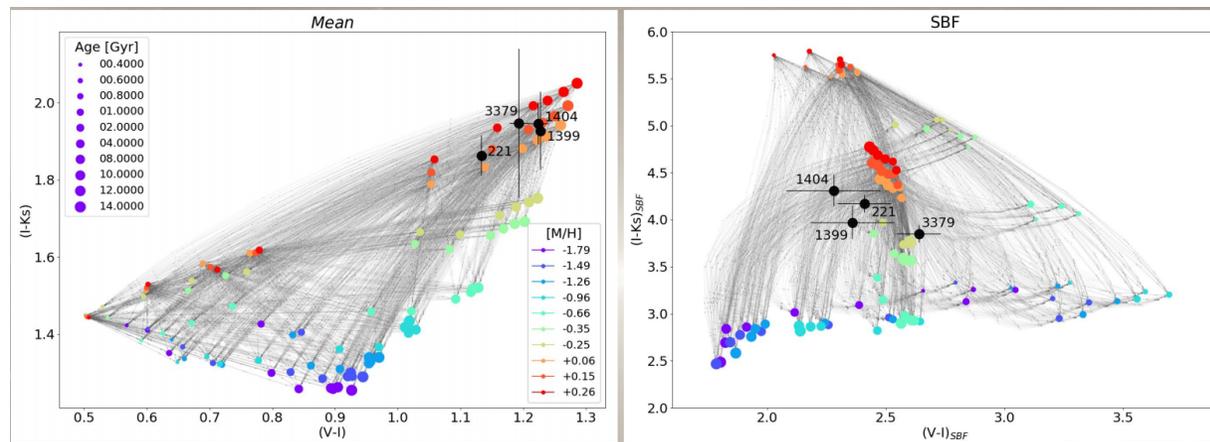
E. Eftekhari: NIR spectral indices

J, H, K bands covered →
galaxy evolution studies in the NIR

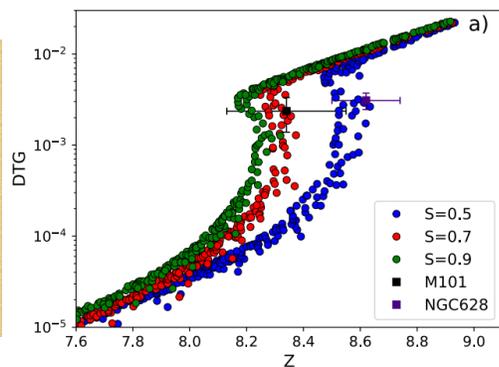
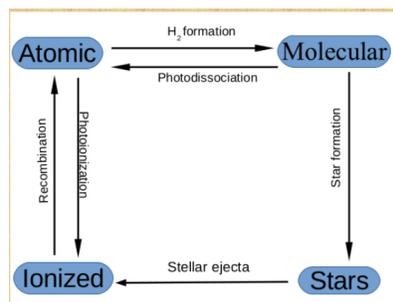


P. Rodriguez Beltran: Surface Brightness Fluctuations

Reveal the metal-poor content in metal-rich ETGs



Multiphase code for galactic chemical evolution:
metals and dust



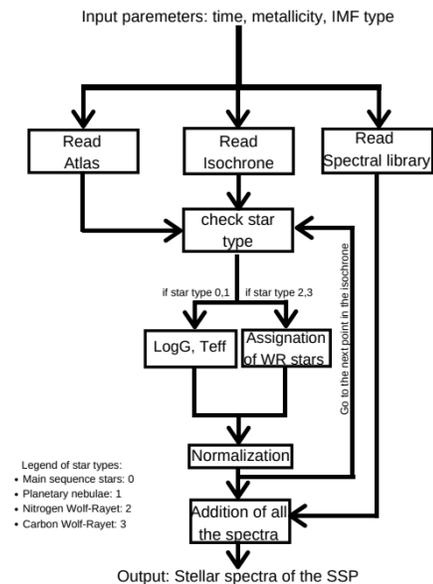
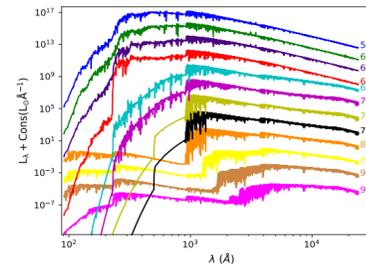
I. Millan Irigoyen:



Models reproduce the DTG-Z relations at high, med Z, but not in the low

HR-pyPopStar

Python-based updated version of PopStar of high resolution SSPs



The nearby Universe: Models & Methods

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**Let's see what new
discoveries lay ahead!**

**Local
Universe
is very
diverse...
and fun!**

