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The aging (early spectral evolution) of SN 2020jgl (Xarel-Io) and the quest for baby supernovae with GTC











SIRAH Supernovae in the InfraRed Avec Hubble

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Type Ia supernova cosmology in the near infrared

Type la supernovae (SNe la) are the most precise SN Ia are more natural standard candles in the near extragalactic distance indicators to date. Empirical infrared (NIR). Both theoretically and empirically, it has been found that NIR SN Ia peak magnitudes relations between peak magnitude, light curve are similar down to ~0.11 mag (in H-band). stretch and color, allow the standardisation of their optical peak magnitudes down to ~0.08mag. Avelino et al. (2019)



 $m_B^{corr} = m_B^{obs} + \alpha x 1 - \beta c$



The main reason for the reduced scatter is dust extinction mitigation at NIR wavelengths.

However, there is a lack of SN Ia observations in the NIR...









HST Cycles 27 and 28 approved program to obtain NIR observations of 24 Hubble-flow (0.02 < z < 0.07) SN la with WFC3/IR [2 objects per month Feb 20-21]

Photometry in 5 NIR bands: F098M, F105W, F125W, F140W, F160W 2 epochs of grism spectroscopy: G102, G141

3 key science goals: \star rest-frame NIR SEDs for WFIRST ★ 2.7% NIR SN la H₀ measurement ***** HST Hubble diagram to z = 0.6 (w/RAISINs; high-z)

Around 2 week turnaround for HST: we need to trigger on very young SN Ia (few days post explosion) want near-max data.

Discovery: building on SN search surveys (ZTF, ATLAS, PS, ASASSN) and public brokers (ALeRCE, Lasair, Anteres) **Classification:** relying on ground-base **ToO programs** with large aperture telescopes (e.g. GTC!)



PI: Saurabh Jha (Rutgers)

erved flux (erg/cm²/s/A)



1.Surveys discover SN candidates



3. SIRAH searches for good young SNIa candidates from brokers





4. ToO programs request spectra to classify the best candidates

5. If a young SNIa, HST is triggered!

first ever rest-frame NIR spectrum of a maximum light SN Ia from space!





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Proposal GTC52-20A (PI: Lluís Galbany): Follow-up of infant SNe within 48h from explosion

10 classifications of SN candidates

SN name	Туре	Time from discovery	
ZTF20aaxvzja - SN2020itj	II	15h 05.84m	
ZTF20aaynrrh - SN2020jfo	П	16h 51.43m	
ATLAS20Iti - SN2020jgl	la	14h 34.39m	
ATLAS20Its - SN2020jhf	la-91bg	1d 11h 48.75m	
ZTF20abaovyz - SN2020kjt	II	18h 58.49m	Tr by
ZTF20abapyxl - SN2020kku	la	18h 28.06m	
ZTF20abbhyxu - SN2020kyx	la	22h 02.15m	
ZTF20abazeye - AT2020krl	CV	3d 17h 44.88m	-
ZTF20abbplei - SN2020lao	IcBL	18h 45.72m	P
ZTF20abisitz - SN2020nny	П	17h 12.89m	

XIV.0 Reunión científica

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 GTC -16.1 d $$
 SALT -15.2 d $$
 GTC -15.1 d
 GTC -14.1 d
 LCOGT -12.6 d $$
 GSP -12.5 d
 GSP -10.6 d
 LCOGT -9.6 d $$
 GSP -8.7 d
 LCOGT -3.8 d $$
 LT -3.2 d
 LT -3.2 d LT -0.2 d
 LT -3.2 d LT -0.2 d GSP 0.3 d
 LT -3.2 d LT -0.2 d GSP 0.3 d UCB 1.1 d
 LT -3.2 d LT -0.2 d GSP 0.3 d UCB 1.1 d GSP 4.3 d
 LT -3.2 d LT -0.2 d GSP 0.3 d UCB 1.1 d GSP 4.3 d UCB 5.1 d
LT -3.2 d LT -0.2 d GSP 0.3 d UCB 1.1 d GSP 4.3 d UCB 5.1 d GSP 11.2 d
LT -3.2 d LT -0.2 d GSP 0.3 d UCB 1.1 d GSP 4.3 d UCB 5.1 d GSP 11.2 d SALT 13.6 d
LT -3.2 d LT -0.2 d GSP 0.3 d UCB 1.1 d GSP 4.3 d UCB 5.1 d GSP 11.2 d SALT 13.6 d GSP 19.2 d





Impact and prospects for the future

- Continuing GTC-OSIRIS program for the classification of very young SNe, and allow further follow up that lead to individual studies of the most interesting objects
- New GTC-EMIR program to complement NIR spectroscopy of SIRAH targets and enable **SNIa** subtyping
- New NOT/TNG program to complement NIR photometry of SIRAH targets to better populate the H-band light curve, and also provide J-band with prospect of constructing a J and H-band independent Hubble diagrams.
- Long-term project: The *Carnegie Supernova Project* (CSP III) will be focused on the study of infant SNe, starting in mid 2021, right after SIRAH observations are finished.
- Enhance the participation of *Spanish researchers/institutions* in SN physics/cosmology international collaborations (SIRAH, CSP, DES, LSST-DESC...)



