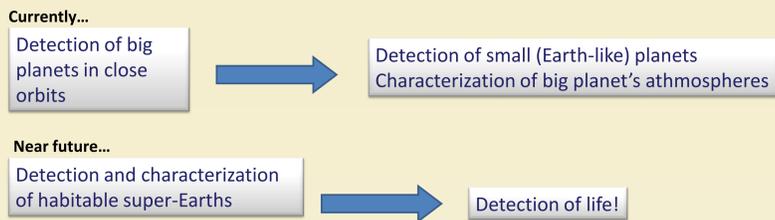
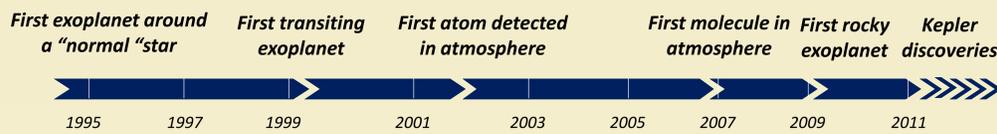
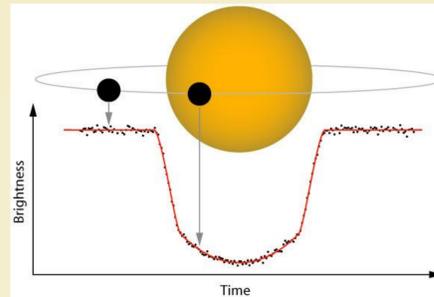


Exoplanet searches using the transit technique are providing new findings at a fast pace. Most exoplanet transit detection programs that are currently underway are focused on large catalogs of stars with no pre-selection. This necessarily makes such surveys quite inefficient, because large amounts of data are processed for a relatively low transiting planet yield. In this work we investigate a method to increase the efficiency of a targeted exoplanet search with the transit technique by preselecting a subset of candidates from large catalogs of stars. Assuming spin-orbit alignment, this can be done by considering stars that have higher probability to be oriented nearly equator-on (inclination close to 90°). We use activity-rotation velocity relations for low-mass stars to study the dependence of the position in the activity- $v\sin i$ diagram on the stellar axis inclination. We compose a catalog of G-, K-, and M-type main sequence simulated stars using isochrones, an isotropic inclination distribution and empirical relations to obtain their rotation periods and activity indexes. Then the activity- $v\sin i$ diagram is filled and statistics are applied to trace the areas containing the higher ratio of stars with inclinations above 80°. A similar statistics is applied to stars from real catalogs with $\log(R'_{HK})$ and $v\sin i$ data to find their probability of being equator-on. We present the method used to generate the simulated star catalog and the subsequent statistics to find the highly inclined stars from real catalogs using the activity- $v\sin i$ diagram. Several catalogs from the literature are analyzed and a subsample of stars with the highest probability of being equator-on is presented. Assuming spin-orbit alignment, the efficiency of an exoplanet transit search in the resulting subsample of probably highly inclined stars is estimated to be two to three times higher than with a global search considering no pre-selection.

Exoplanets: a recent field

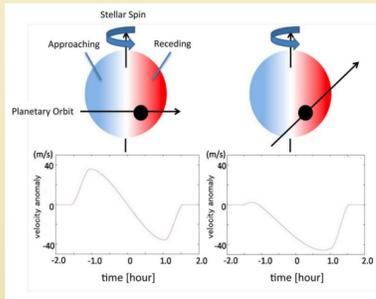


Exoplanet transits

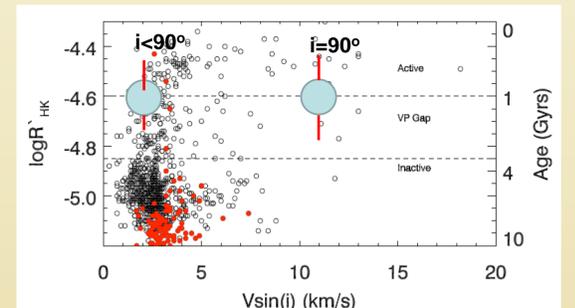
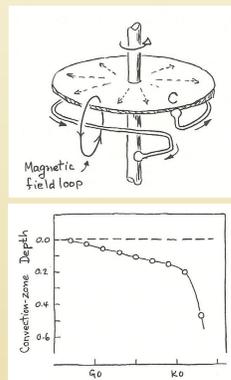


OUR AIM is to contribute to the transit detection efforts by selecting stars being equator-on oriented

Spin-orbit orientation

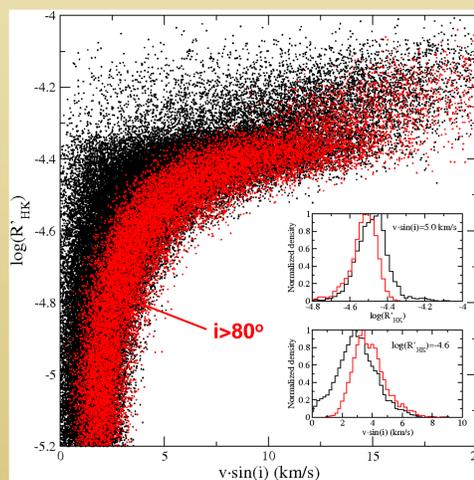
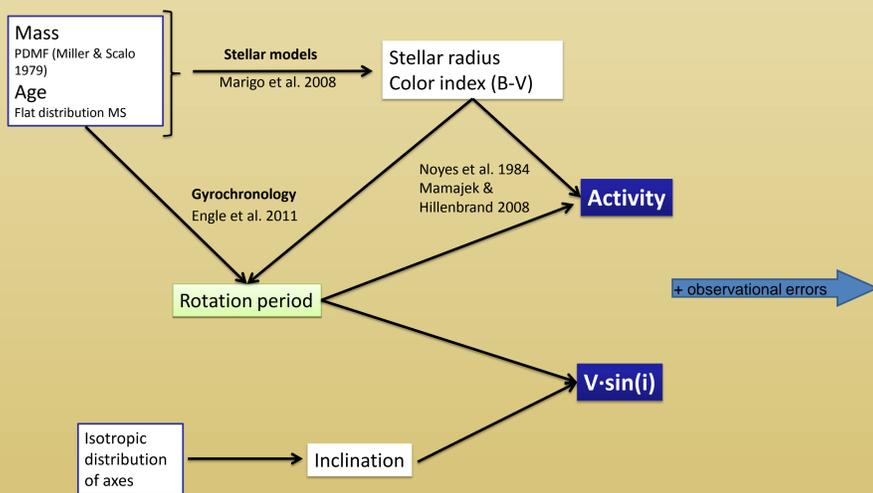


Activity - rotation relations for low mass stars

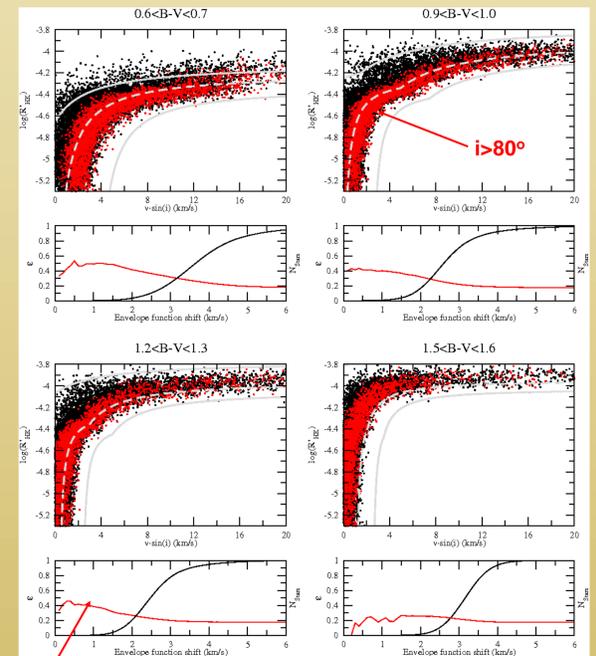


Simulated samples of GKM dwarfs

Simulating Activity & Vsin i



Stars with high inclination are in the red envelope region at the right hand of the total distribution. For a given $\log(R'_{HK})$, stars are expected to have a very similar equatorial velocity in terms of the rotation and activity evolution assumptions, and so $v\sin i$ are mainly caused by different axis projections.



The preselection of stars with high probability of having high inclination axis enhances the success rate of an exoplanet transit search by a factor 2-3.

$$\epsilon_\alpha = \frac{N_{i > \alpha}}{N_T}$$

Efficiency = probability for a star in a given location of the diagram to have an inclination angle above α .
 $\epsilon = 0.173$ for random distribution of orientations.



Related papers

Brown, D. J. A., et al., MNRAS, 423, 1503
 Herrero, E., Ribas, I., Jordi, C., Guinan, E.F., Engle, S.G., 2012, A&A 537, A147

Mamajek, E. E., Hillenbrand, L. A., 2008, ApJ, 647, 1264
 Marigo, P., et al, 2008, A&A, 482, 883
 Noyes, R. W., Hartmann, L. W., Baliunas, S. L., et al. 1984, ApJ, 279, 763